



# SDC Findings and Interpretations

Alex Doner & SDC TEAM

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- 7) Edwin Bernardoni
- 6) Marcus Piquette
- 5) Jamey Szalay
- 4) Andrew Poppe
- 3) Dong Han
- 2) David James
- 1) Colin Mitchell

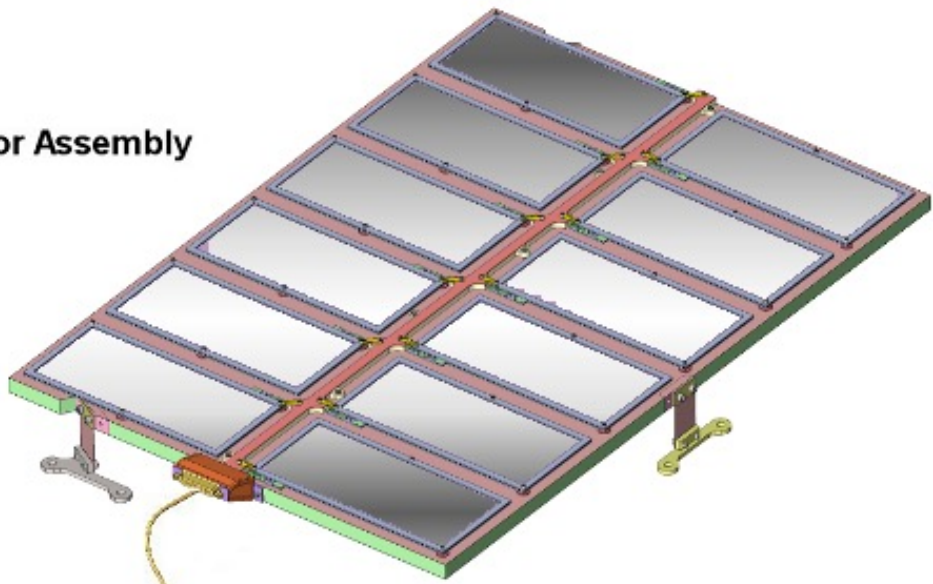




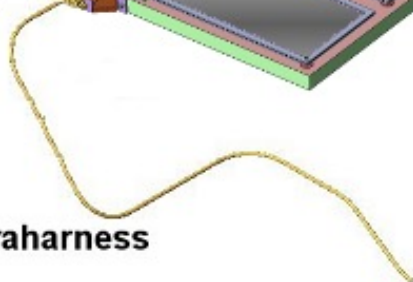
# Student Dust Counter (SDC): Instrument Design



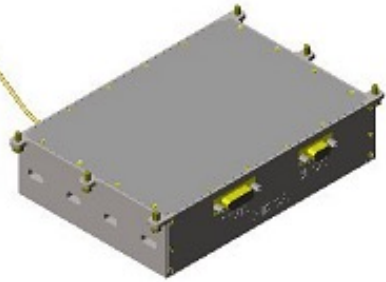
**Detector Assembly**



**Intraharness**



**Electronics Box**

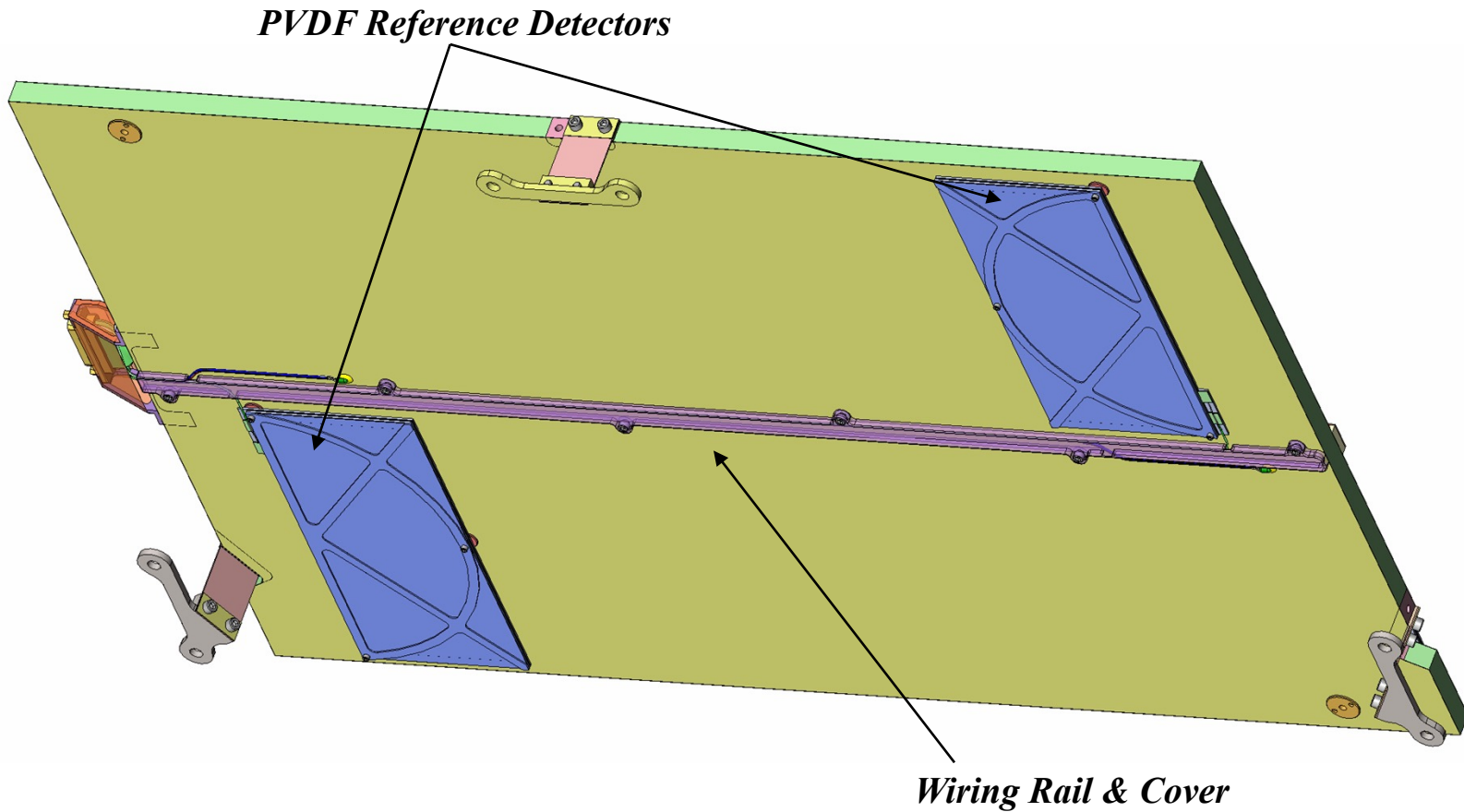


## Key Properties

<b>Active area</b>	<b>0.1 m<sup>2</sup></b>
<b>Mass</b>	<b>1.6 kg</b>
<b>Power</b>	<b>5 w</b>



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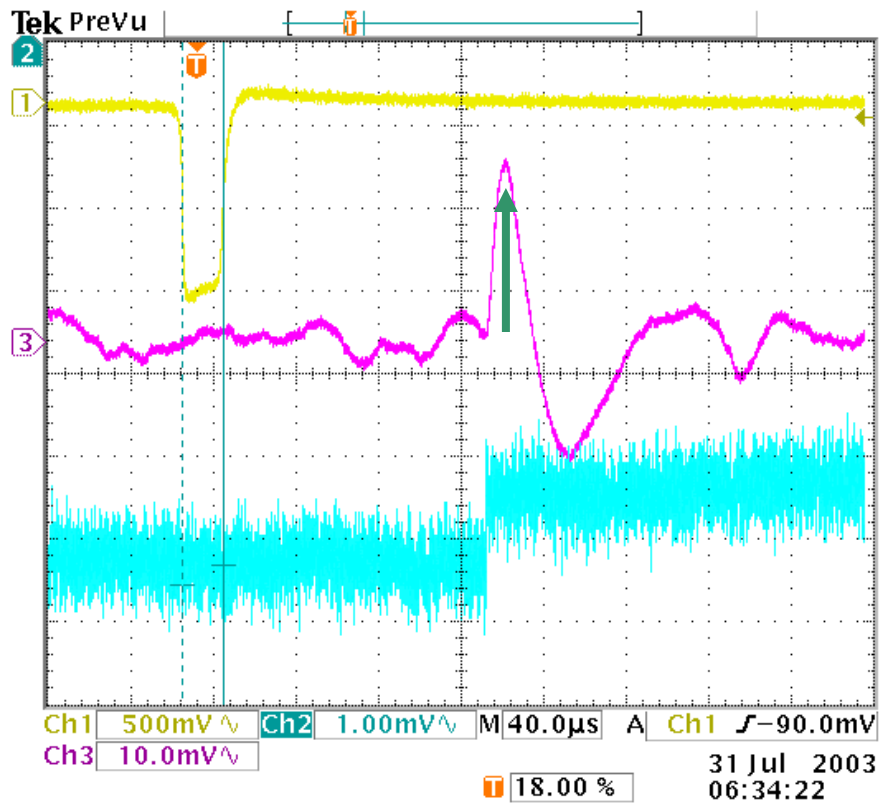
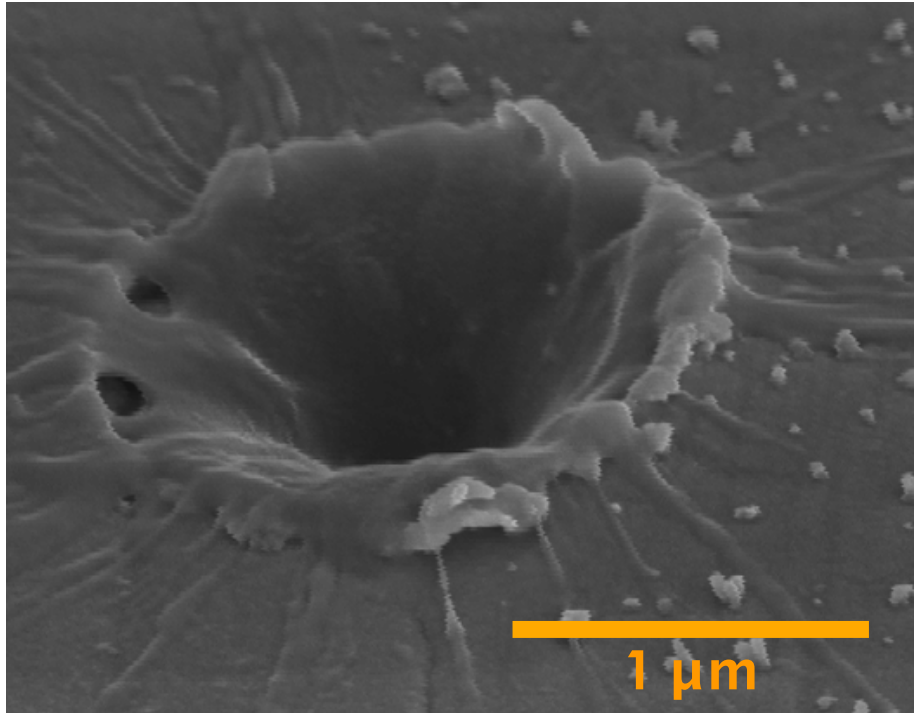


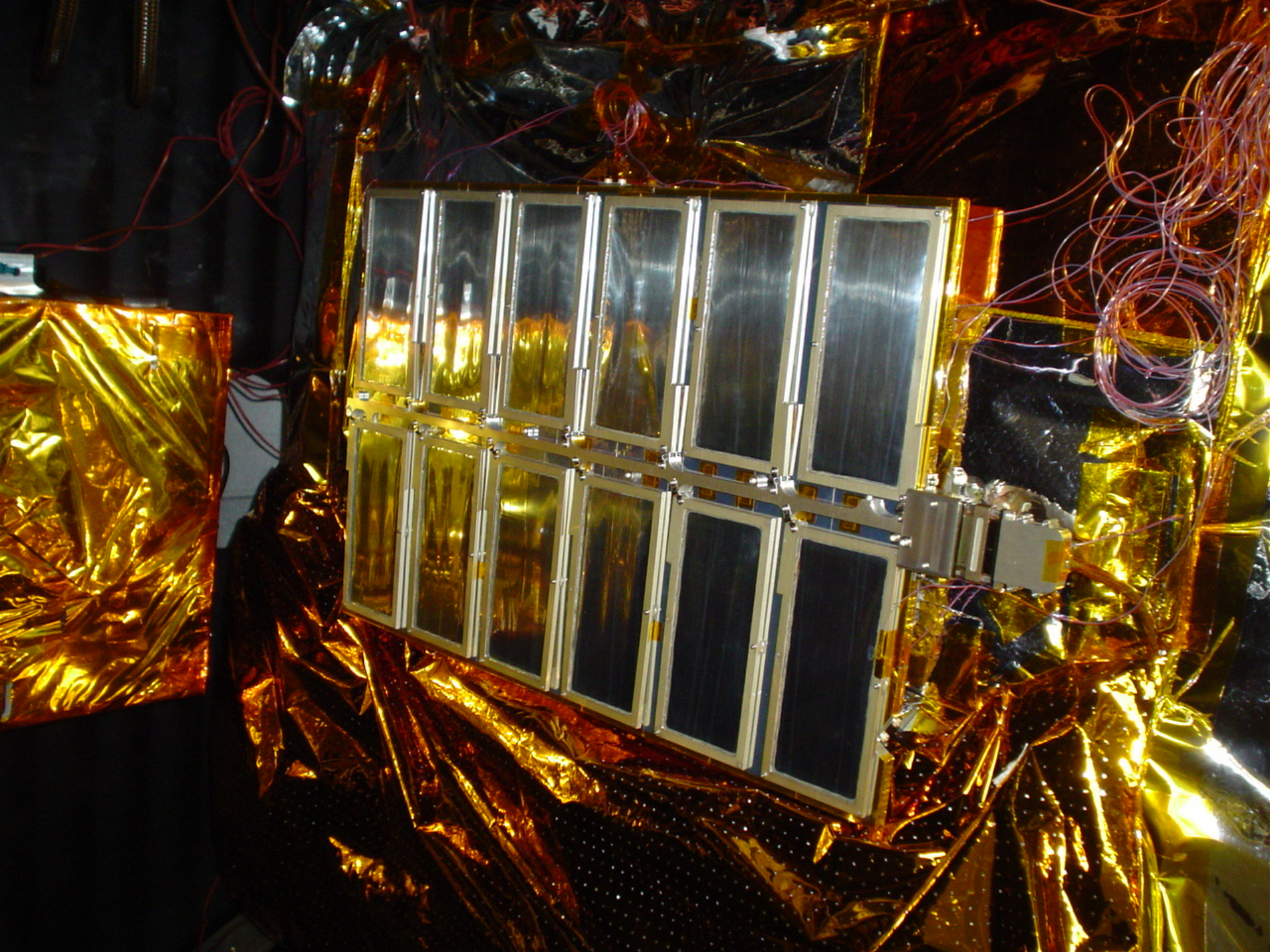


# PVDF Principle of Operation



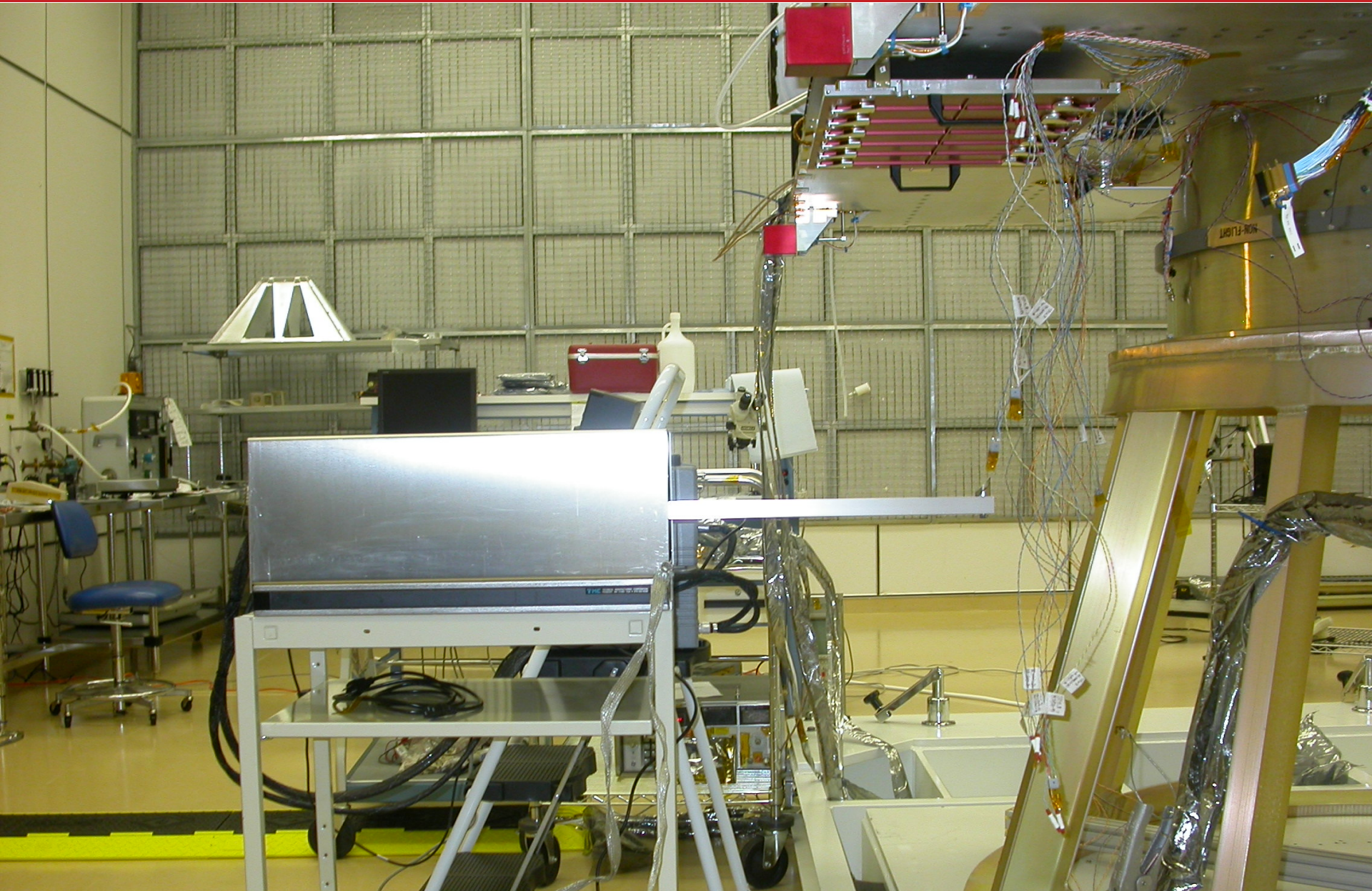
$$N_e = C m^{1.3} v^{3.0}$$





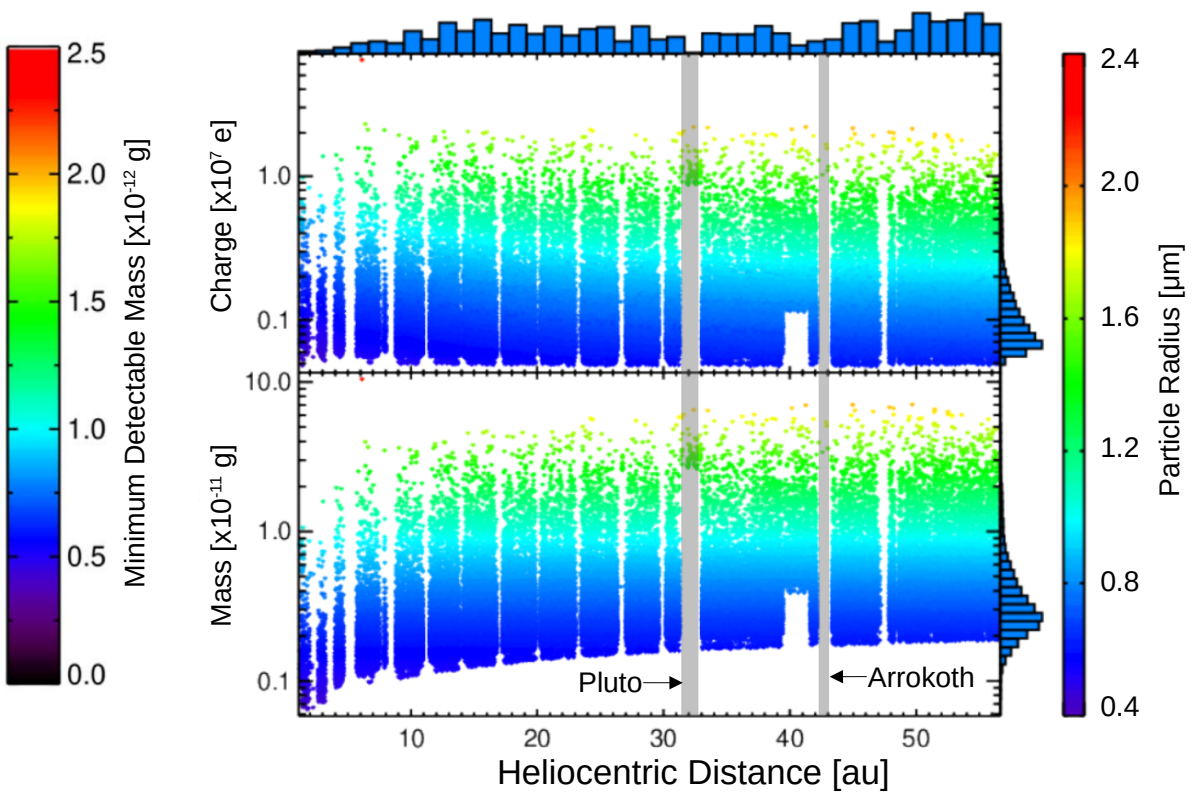
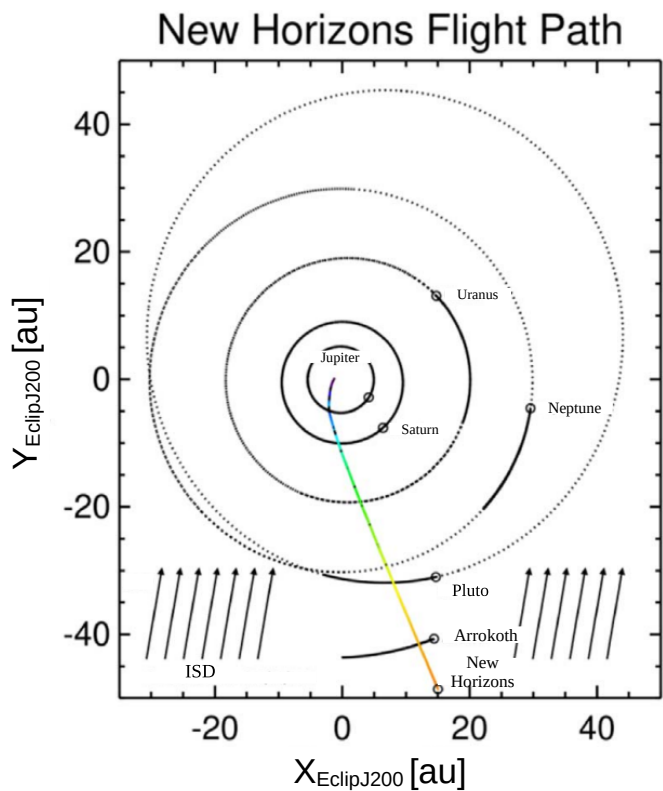


# PVDF Test Setup APL 11/2004





# Data Collection

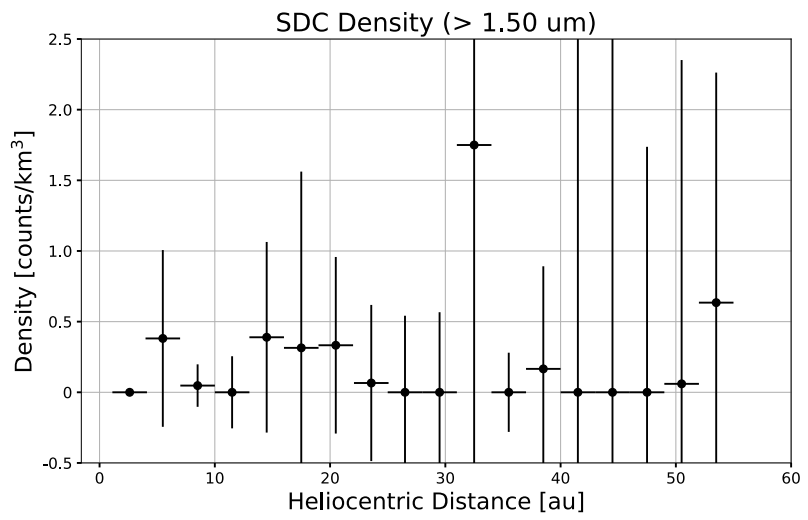
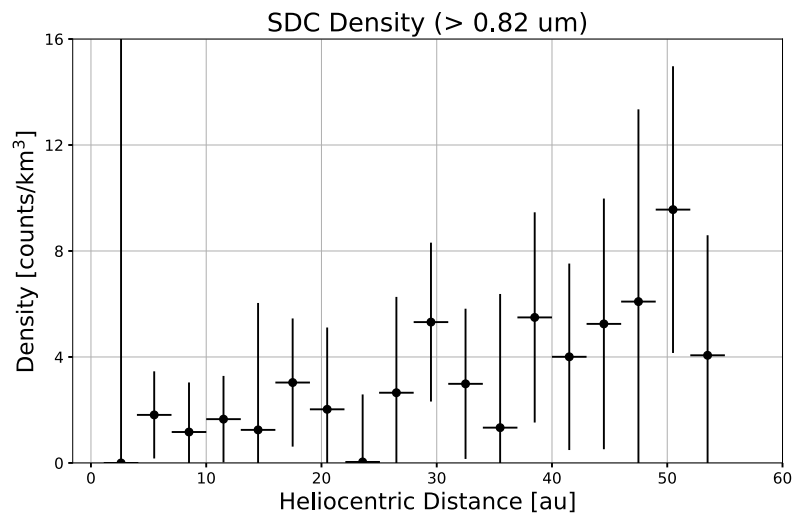
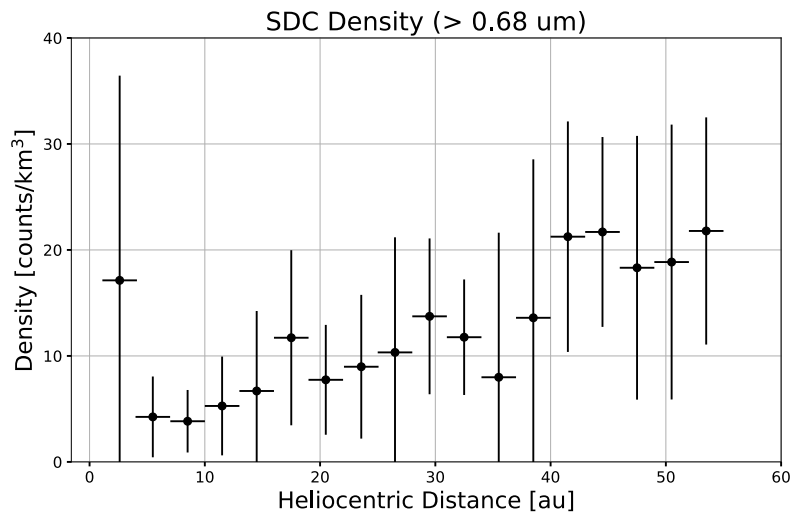
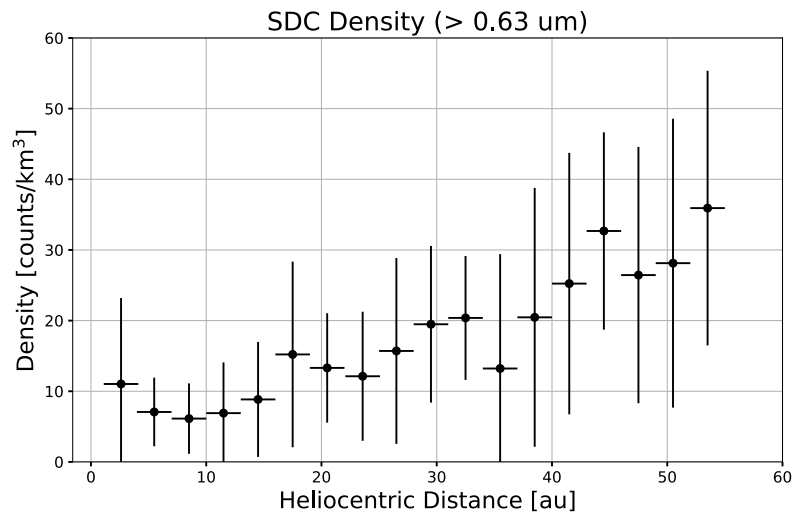




# New Horizons Venetia Burney Student Dust Counter Observes Higher than Expected Fluxes Approaching 60 AU



ALEX DONER <sup>1,2</sup> MIHÁLY HORÁNYI <sup>1,2</sup> FRAN BAGENAL <sup>1</sup> PONTUS BRANDT <sup>3</sup> WILL GRUNDY <sup>4</sup> CAREY LISSE <sup>3</sup>  
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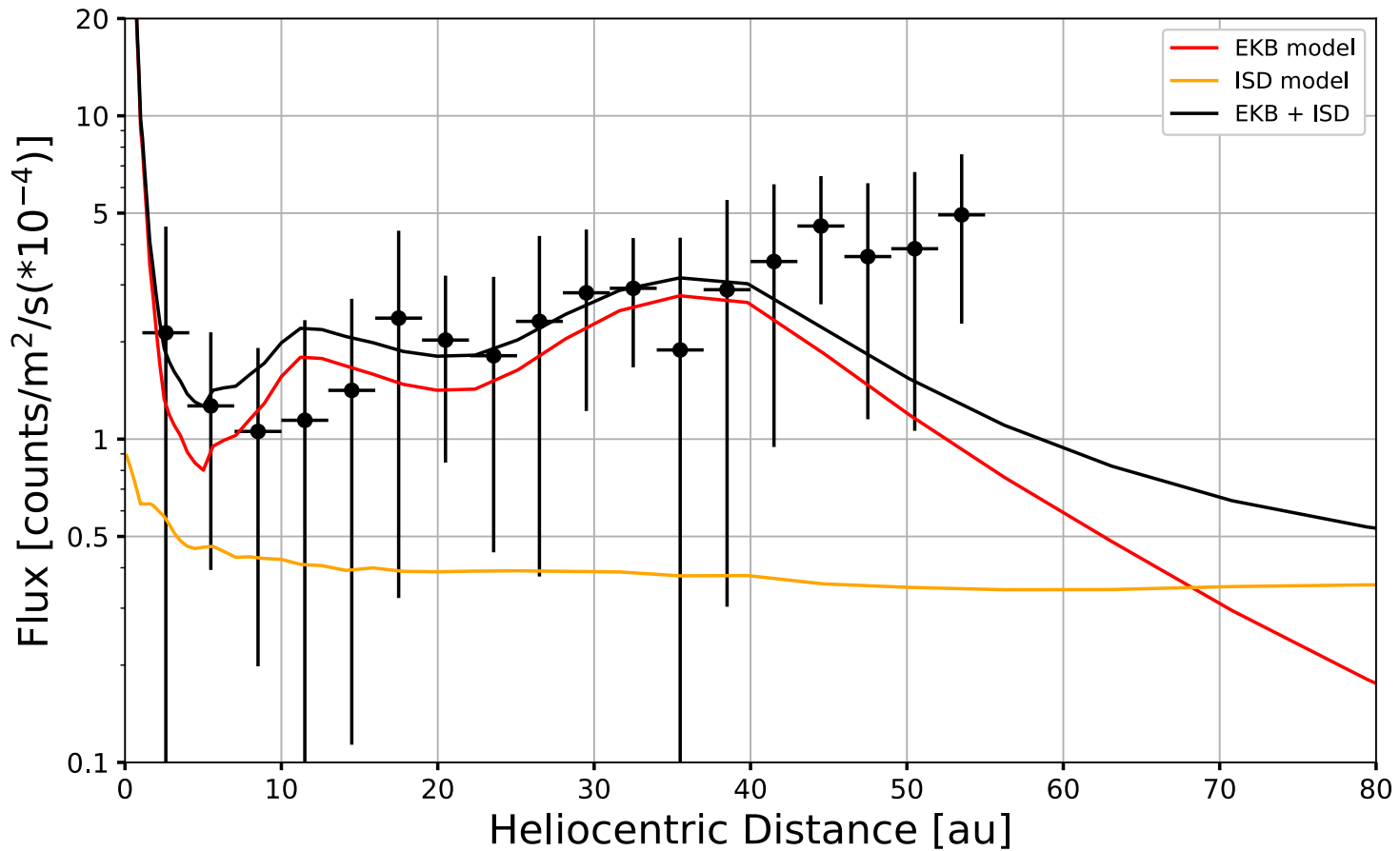






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# SDC Measured Flux (>0.63 μm)





# Radiation Pressure

- Dust is generated by collisions and ISD bombardment.
- Radiation pressure alters the orbital elements.

$\beta = \text{radiation pressure} / \text{solar gravity}$

$$E^* = \frac{1}{2}(\dot{r}^2 + r^2\dot{\theta}^2) - \frac{(1 - \beta)\mu}{r} = E + \frac{\beta\mu}{r},$$

$$J^* = r^2\dot{\theta} = J,$$

$$a^* = \frac{(1 - \beta)ar}{r - 2\beta a} \quad e^* = \sqrt{1 - \frac{(1 - e^2)(r - 2\beta a)}{(1 - \beta)^2 r}}$$



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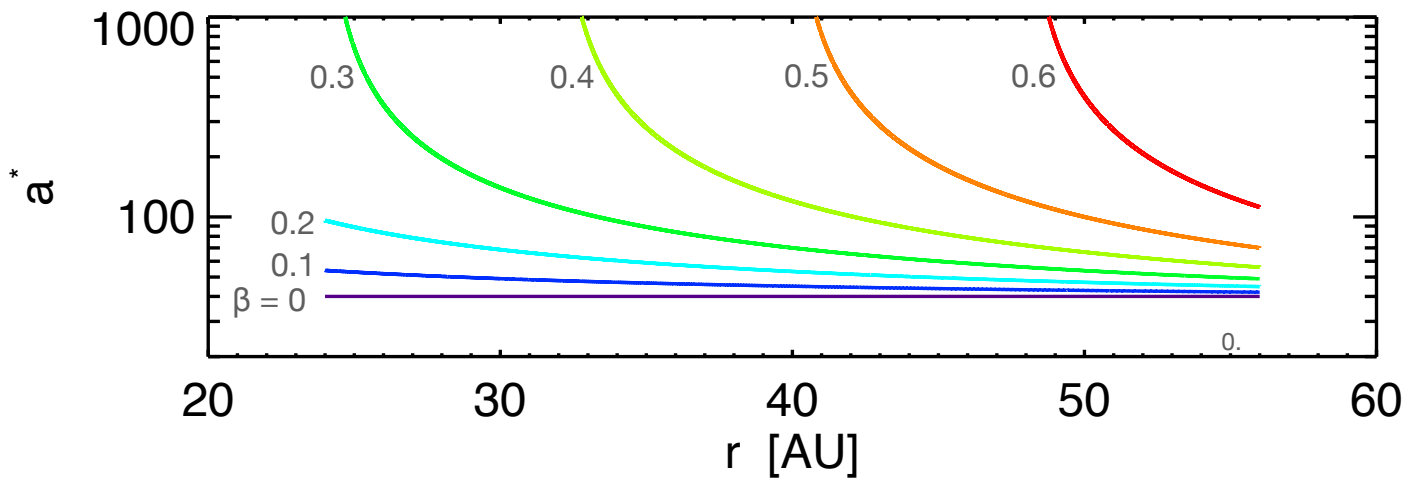
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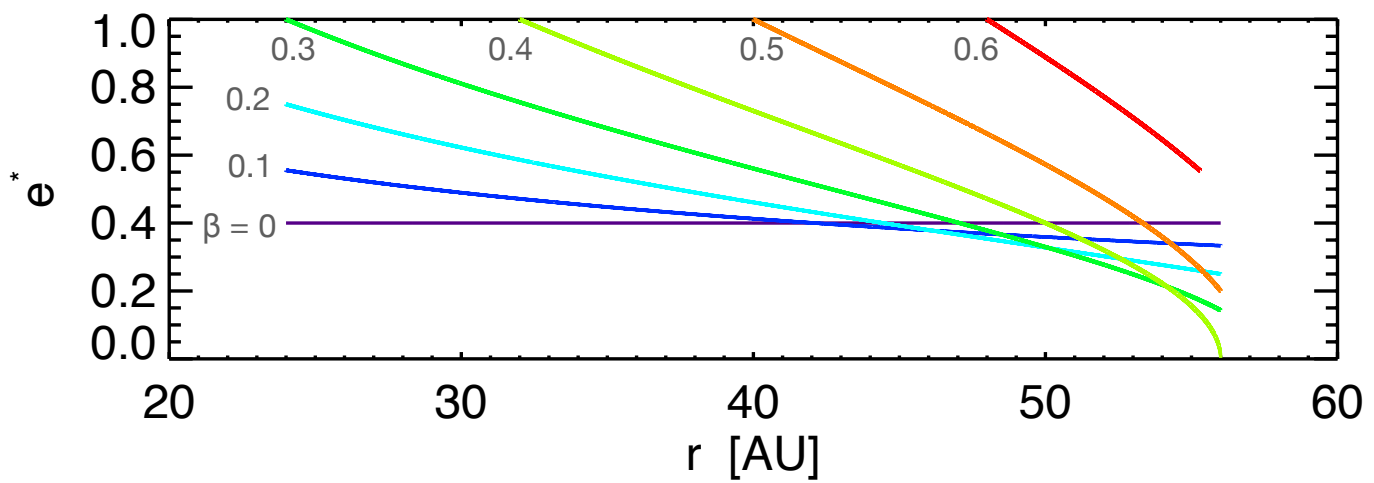
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# Radiation Pressure



Source  
 $a_0 = 40$   
 $e_0 = 0.4$



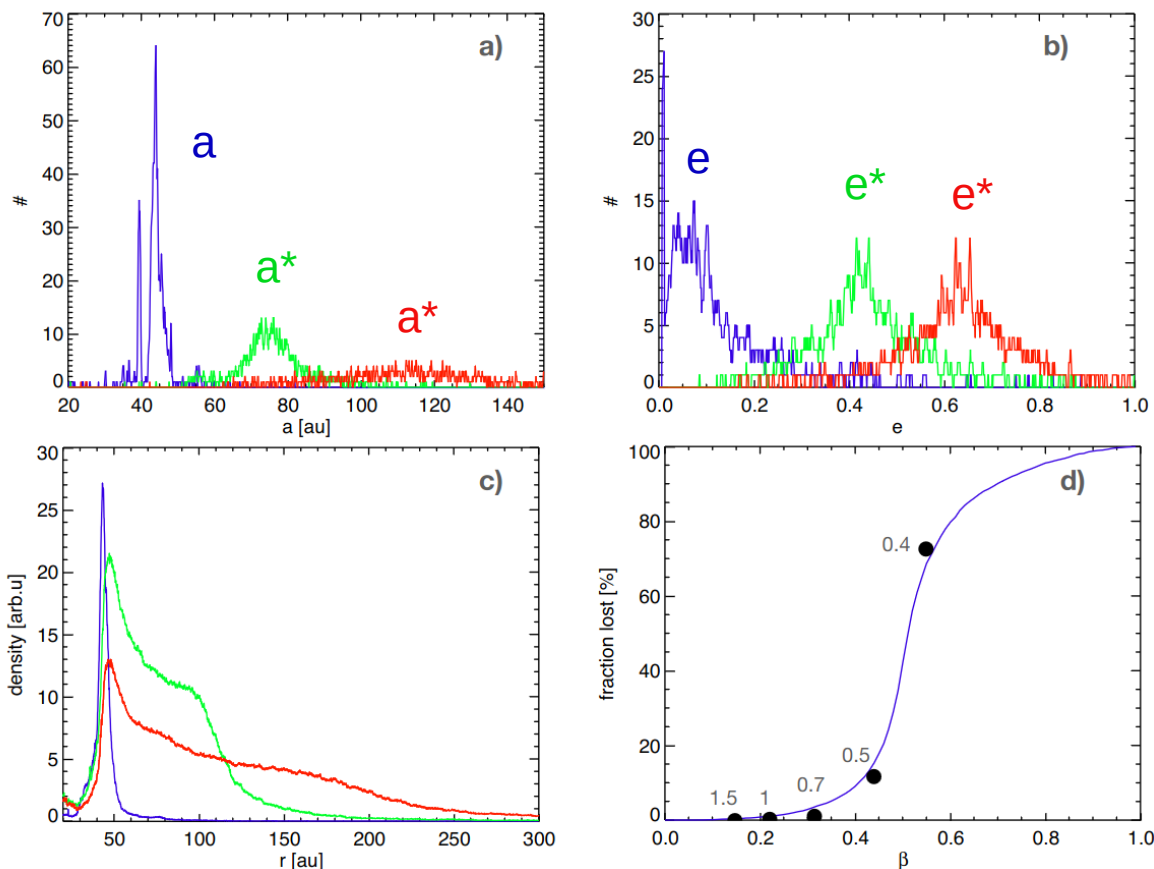


# Initial Dust Density Distribution



Orbital parameter distribution of dust producing bodies and representative dust grains

KBO Source Bodies  
Created Dust:  $\beta=0.3$   
Created Dust:  $\beta=0.4$





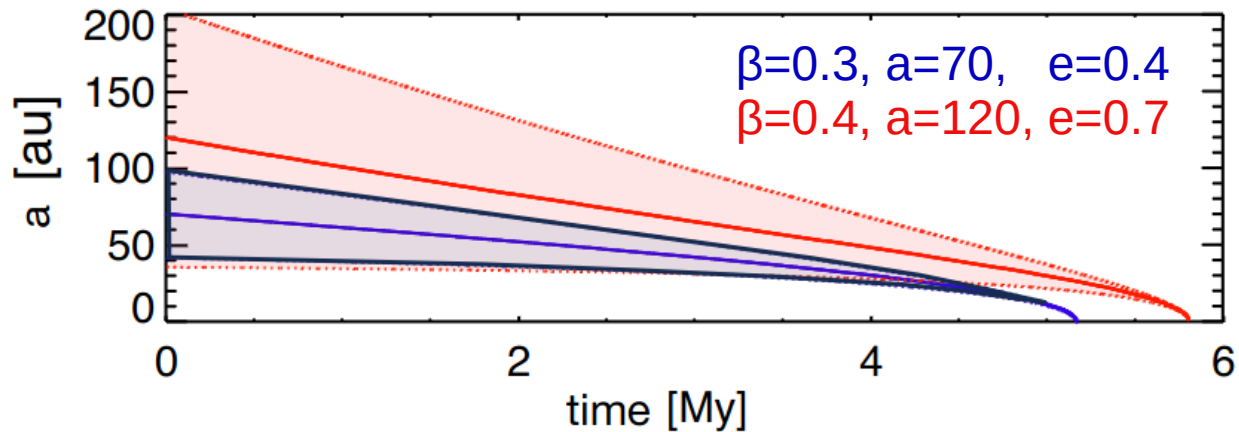
# Orbital Evolution



## Poynting-Robertson Drag

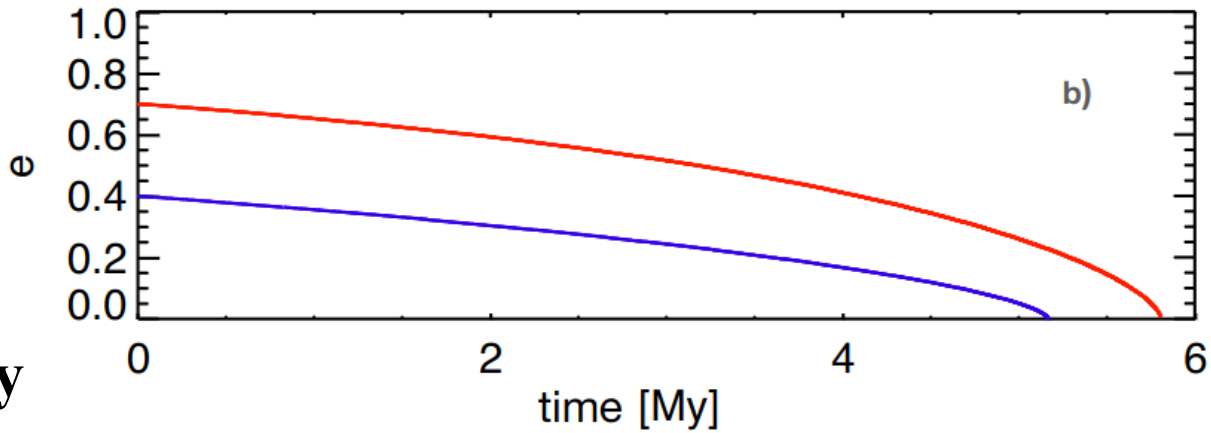
$$\frac{da}{dt} = -c \frac{\beta}{a} \frac{(2 + 3e^2)}{(1 - e^2)^{3/2}}$$

$$\frac{de}{dt} = -c \frac{\beta}{a^2} \frac{5e}{2(1 - e^2)^{1/2}}$$



Particles slowly circularize as they migrate to the Sun.

Typical lifetime ~ 5 My



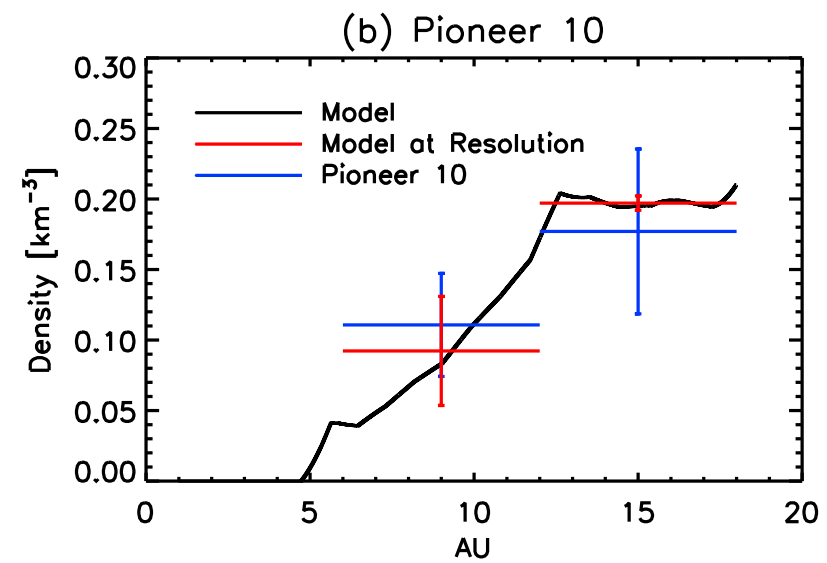
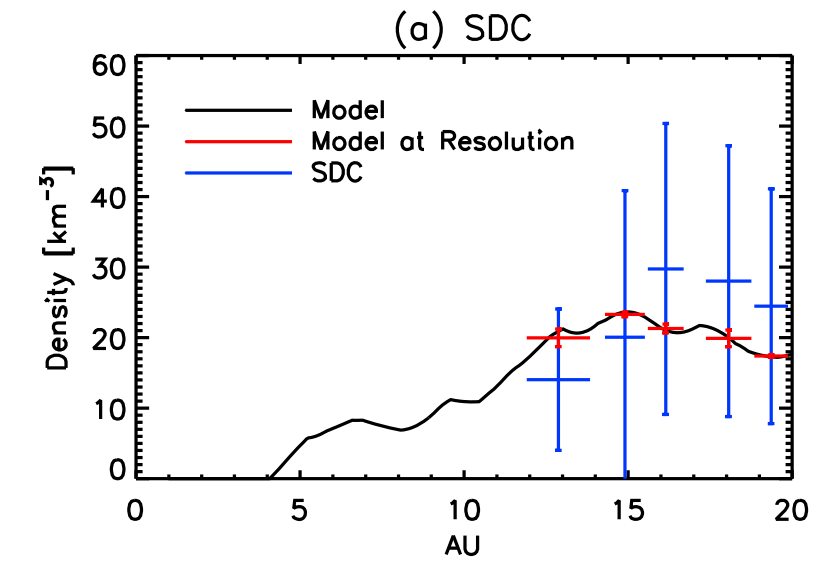
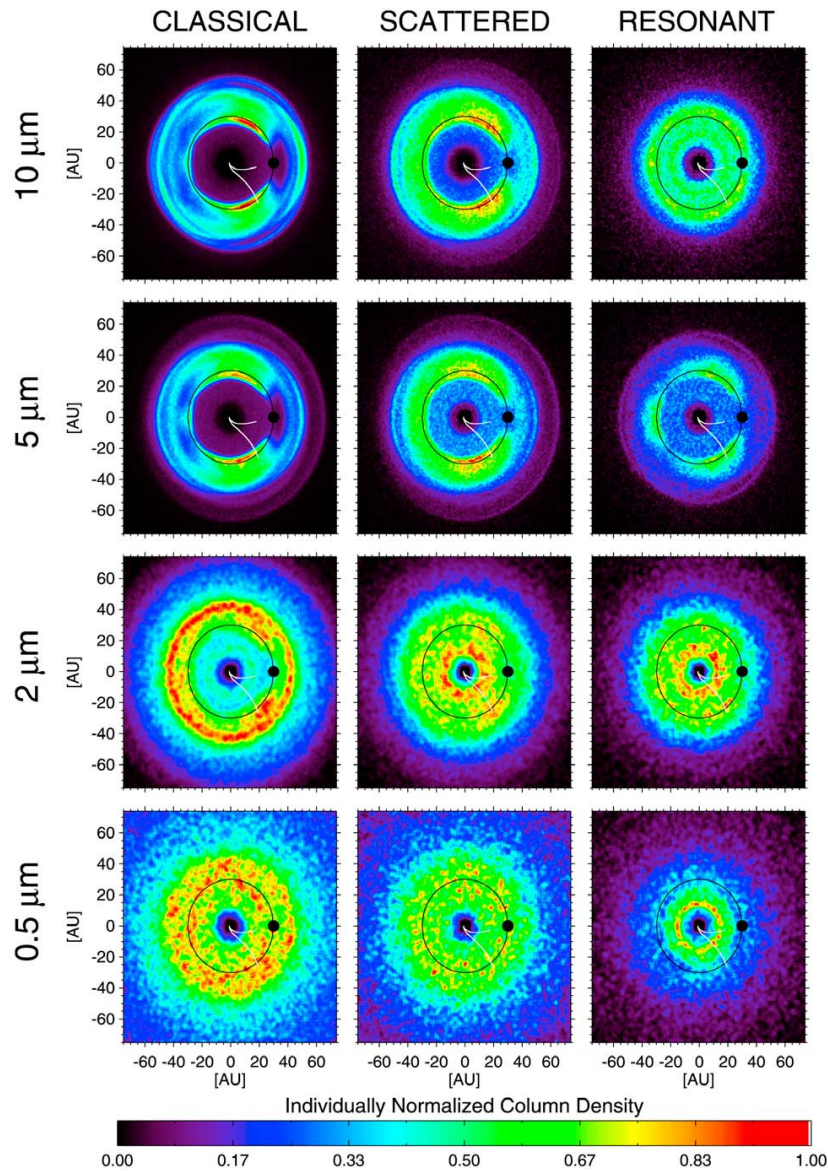


# Constraints on dust production in the Edgeworth-Kuiper Belt from Pioneer 10 and New Horizons measurements

Dong Han,<sup>1</sup> Andrew R. Poppe,<sup>2</sup> Marcus Piquette,<sup>1</sup> Eberhard Grün,<sup>1</sup> and Mihály Horányi<sup>1,3</sup>

Received 24 October 2011; revised 10 November 2011; accepted 13 November 2011; published 28 December 2011.

2011



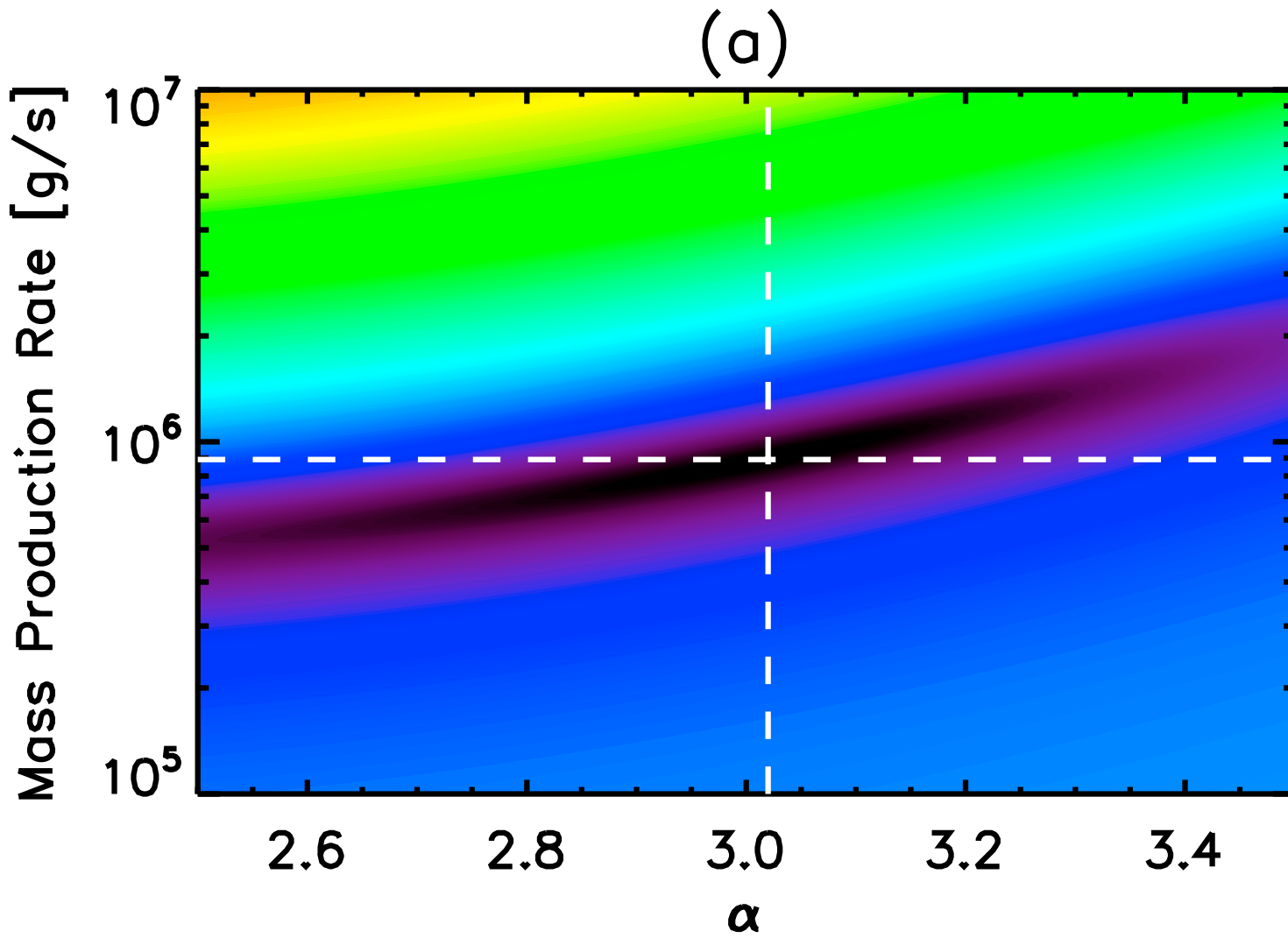


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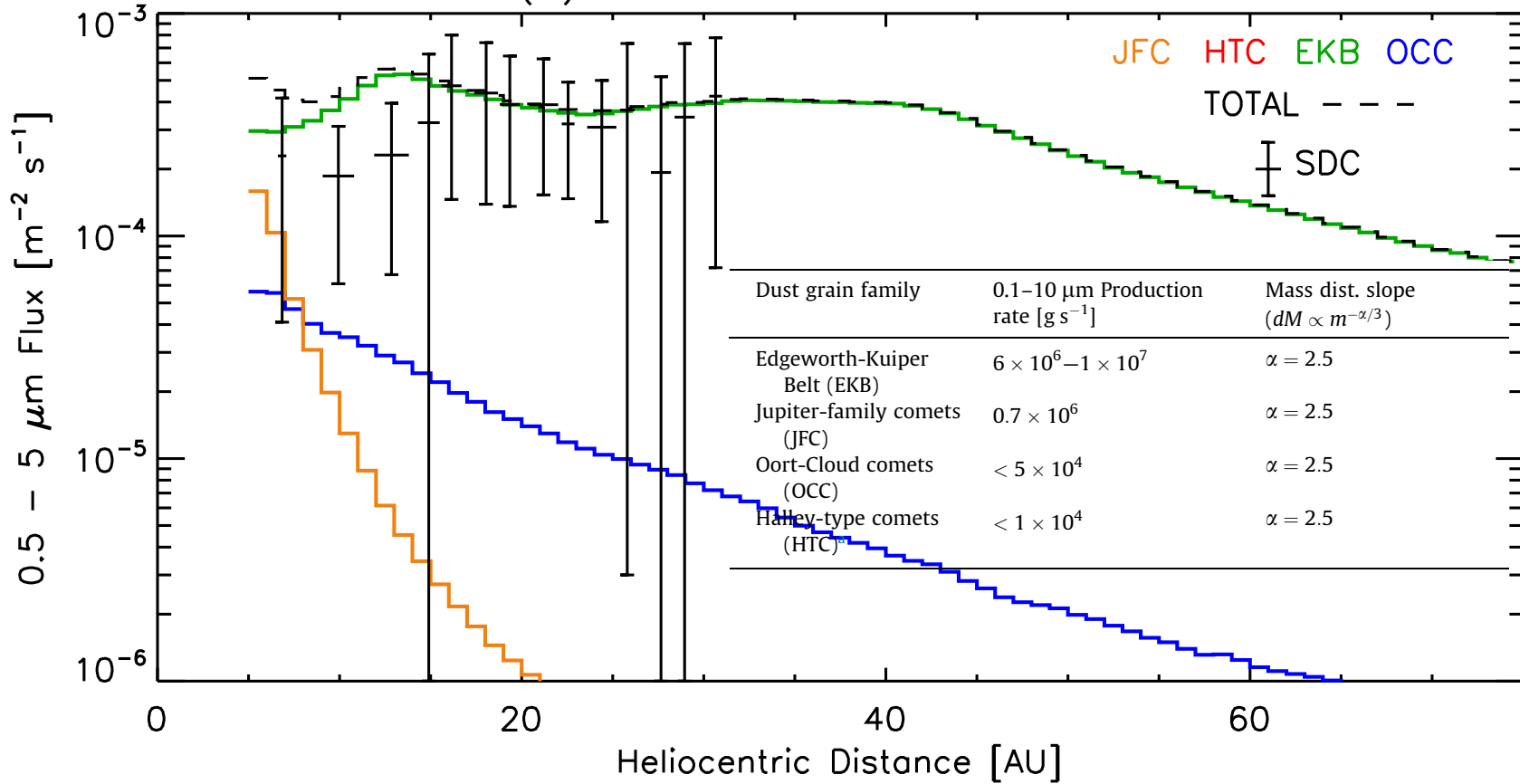
2011







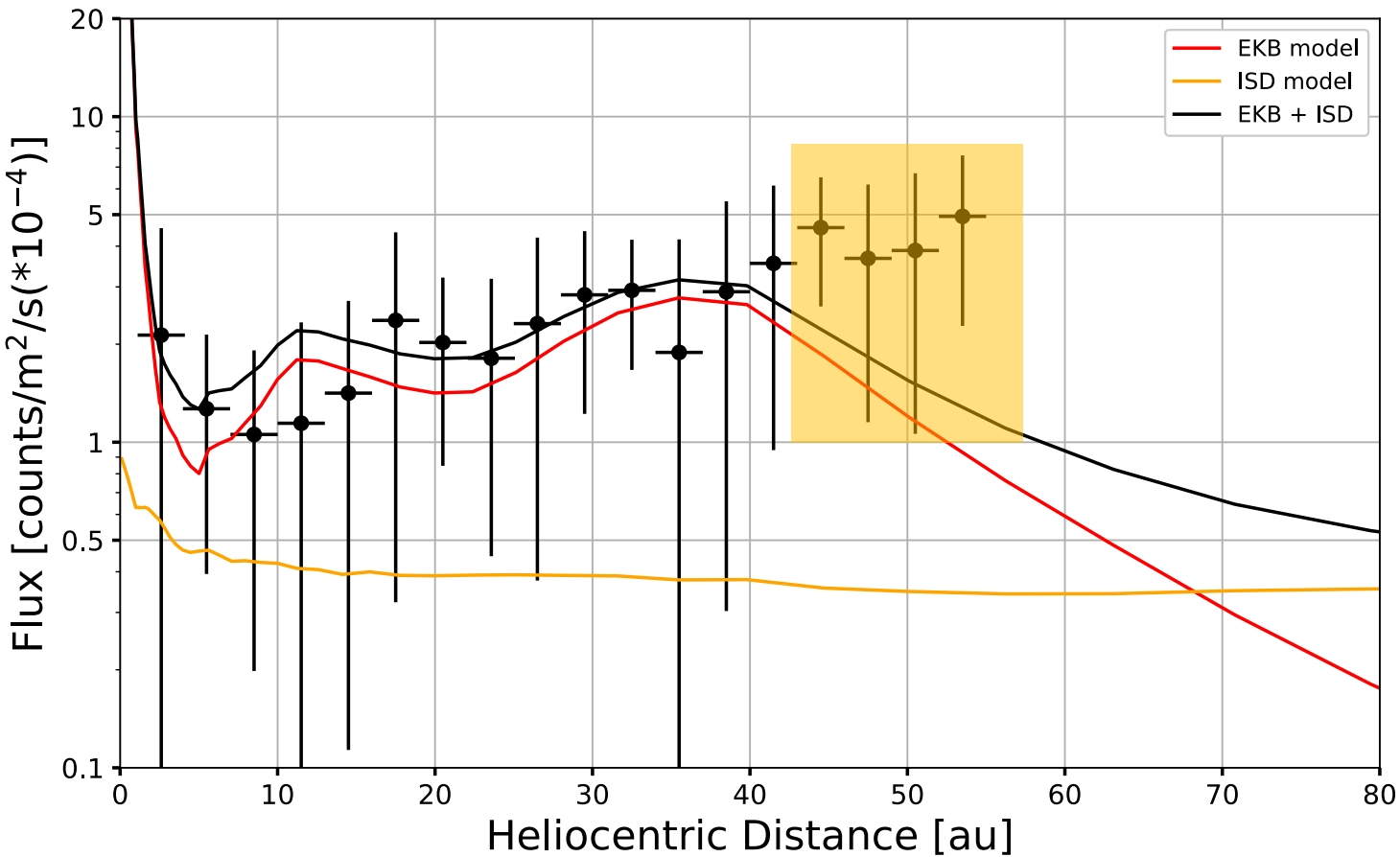
(b) Flux to New Horizons SDC





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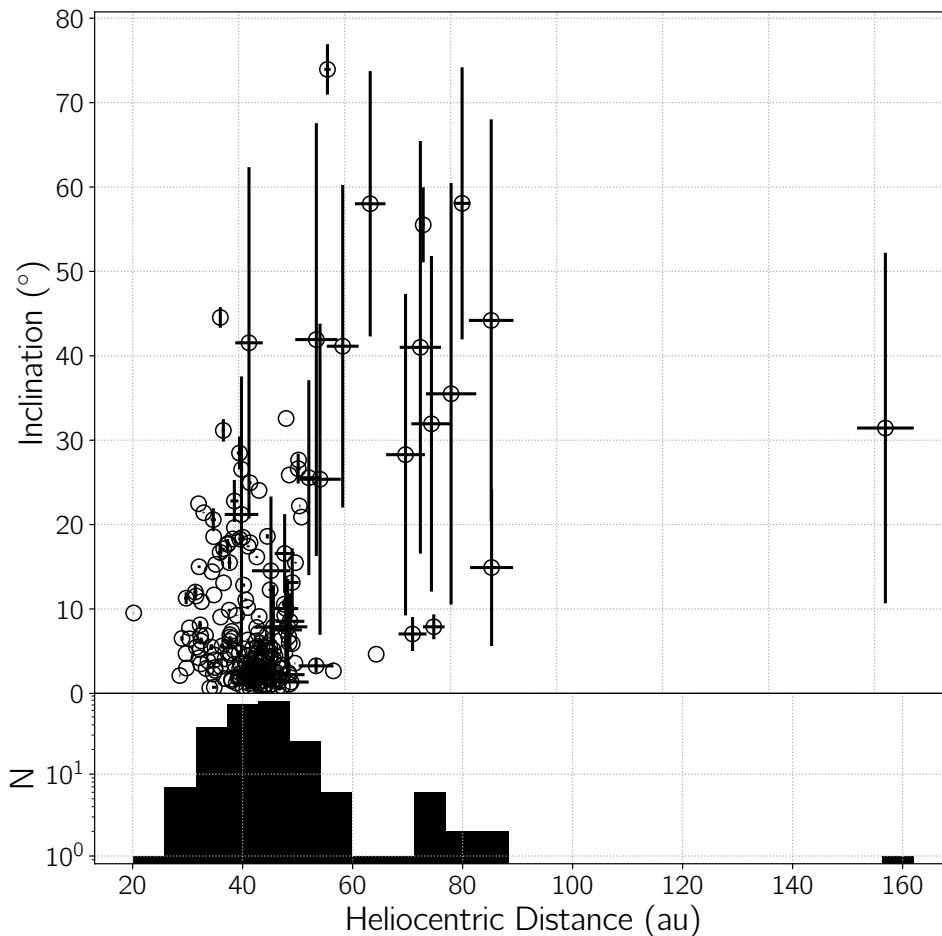




# Extended Kuiper Belt ?

## New Horizons KBO Detections

WESLEY C. FRASER,<sup>1,2</sup> SIMON B. PORTER,<sup>3</sup> LOWELL PELTIER,<sup>1,2,4</sup> JJ KAVELAARS,<sup>1,2</sup>  
ANNE J. VERBISER,<sup>5,3</sup> MARC W. BUIE,<sup>3</sup> S. ALAN STERN,<sup>3</sup> SUSAN D. BENECCI,<sup>6</sup> TAKASHI ITO,<sup>7</sup>  
FUMI YOSHIDA,<sup>8,9</sup> DAVID W. GERDES,<sup>10,11</sup> KEVIN J. NAPIER,<sup>10</sup> HSING WEN LIN,<sup>10</sup>  
SEBASTIEN FABBRO,<sup>1,2</sup> AND HAYDEN SMOTHERMAN<sup>12</sup>





# New Population of KB Dust?



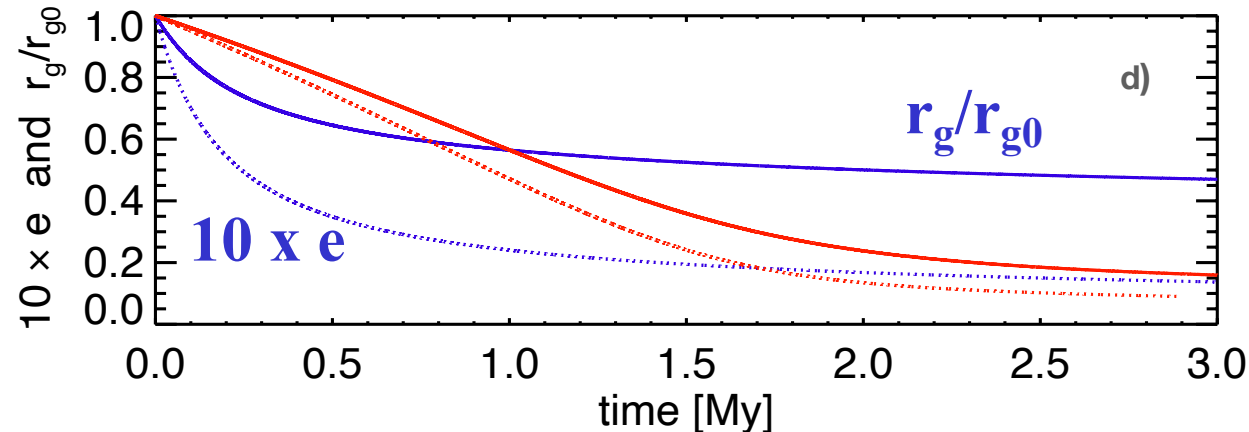
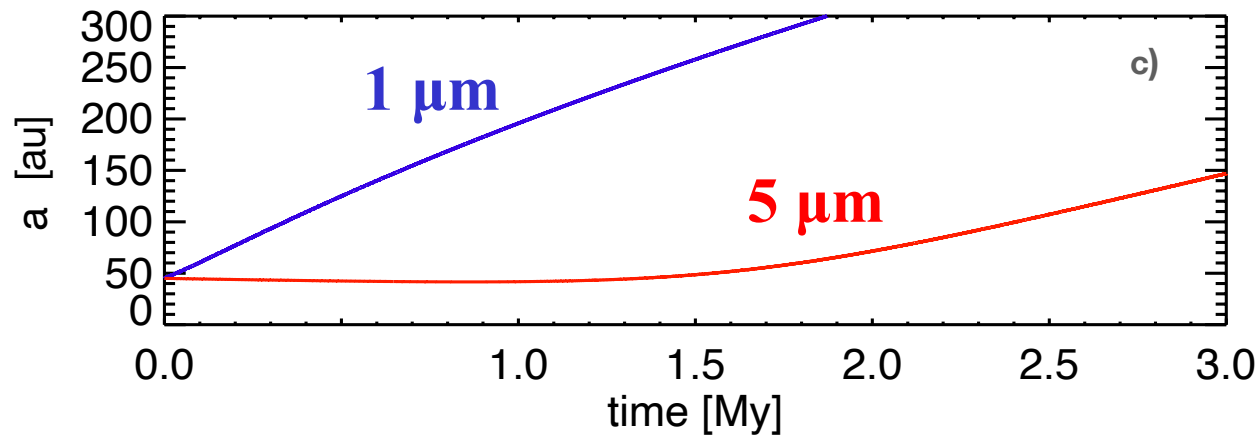
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$$\frac{dr_g}{dt} \simeq 0.4/d^2 \text{ cm/My,}$$

$$\rightarrow + \frac{d\beta}{dt}$$

$$a(t=0) = 45 \text{ \& } e(t=0) = 0.1$$



**Photo-sputtering of ice grains?**

**These never leave the KB.**

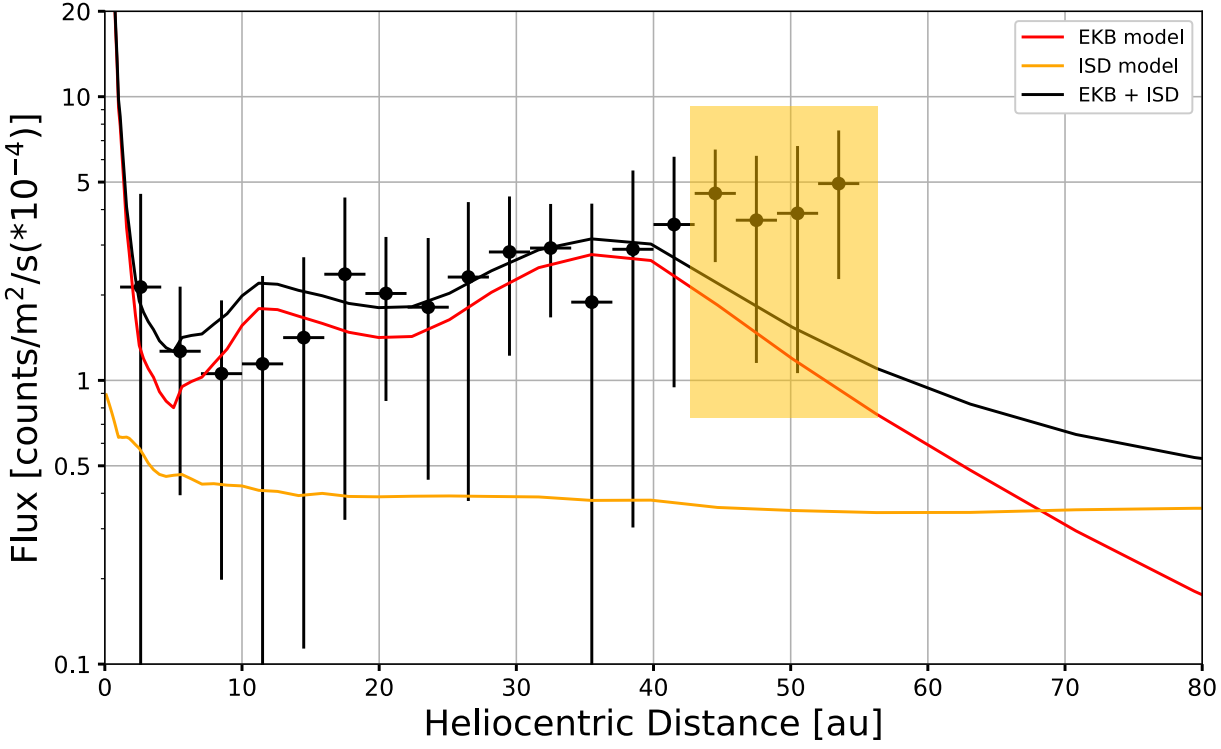


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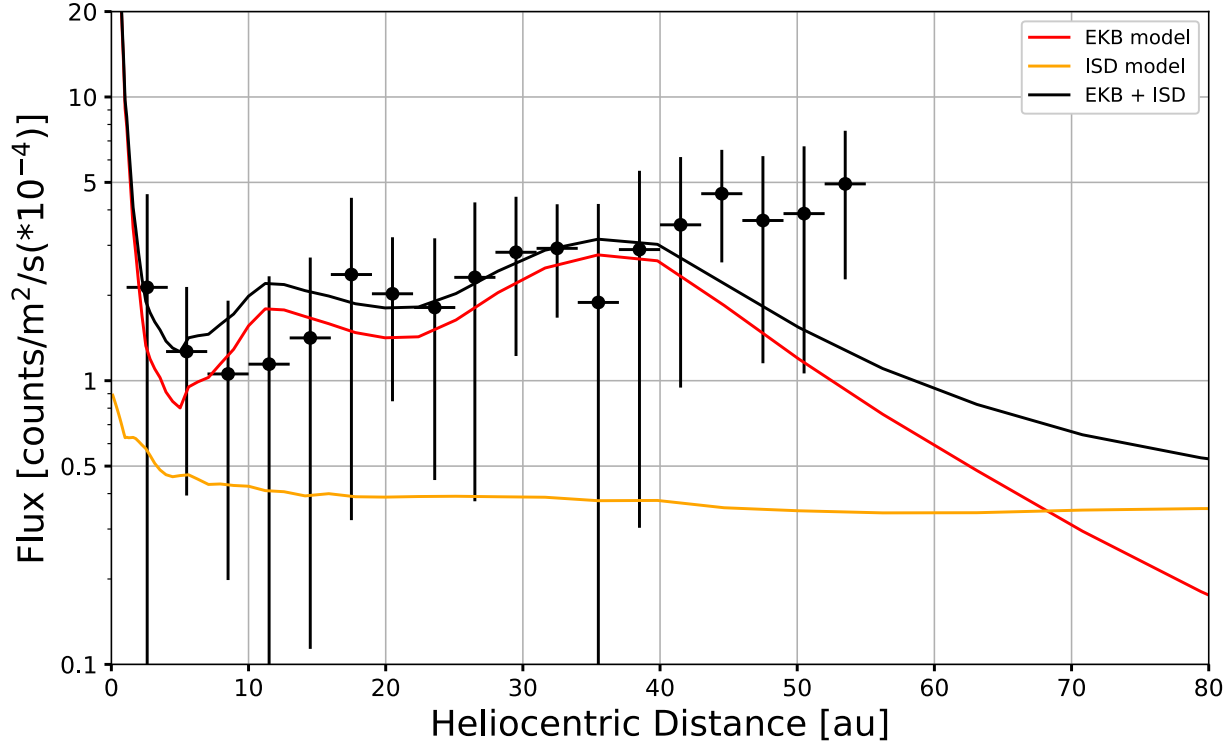
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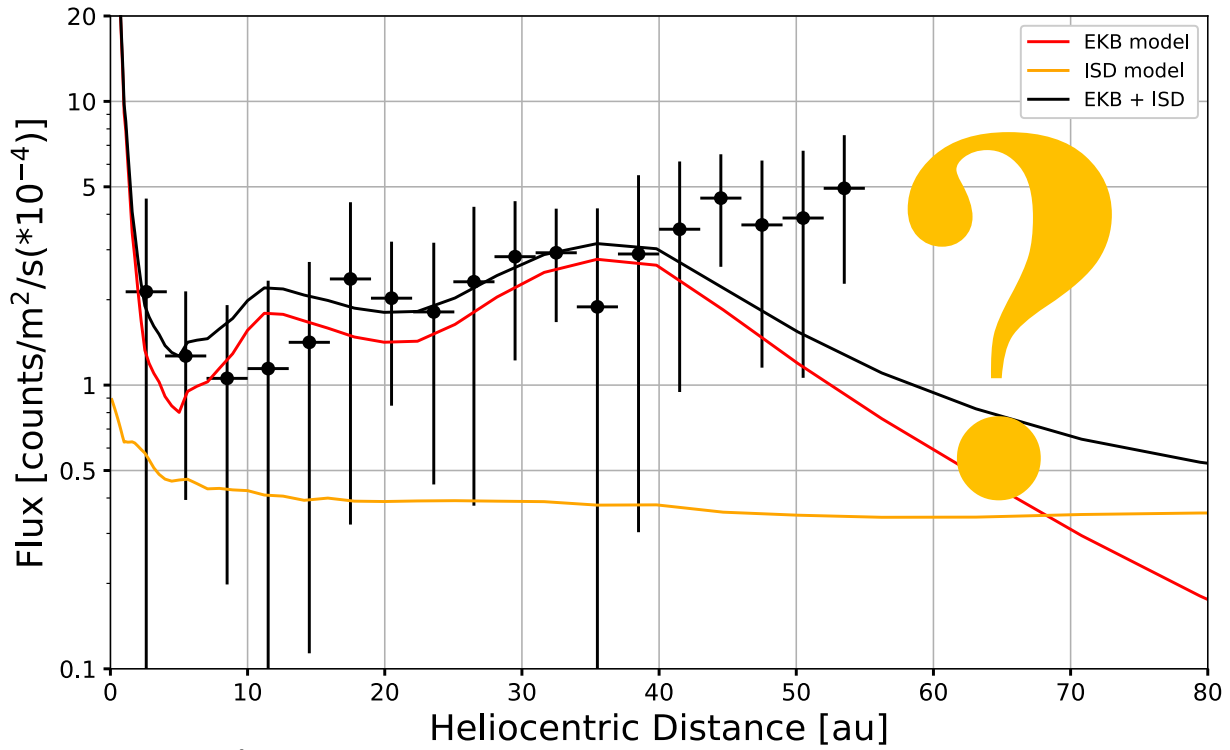


- a) Extended Kuiper Belt
- b) Compositional gradients
- c) .....



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